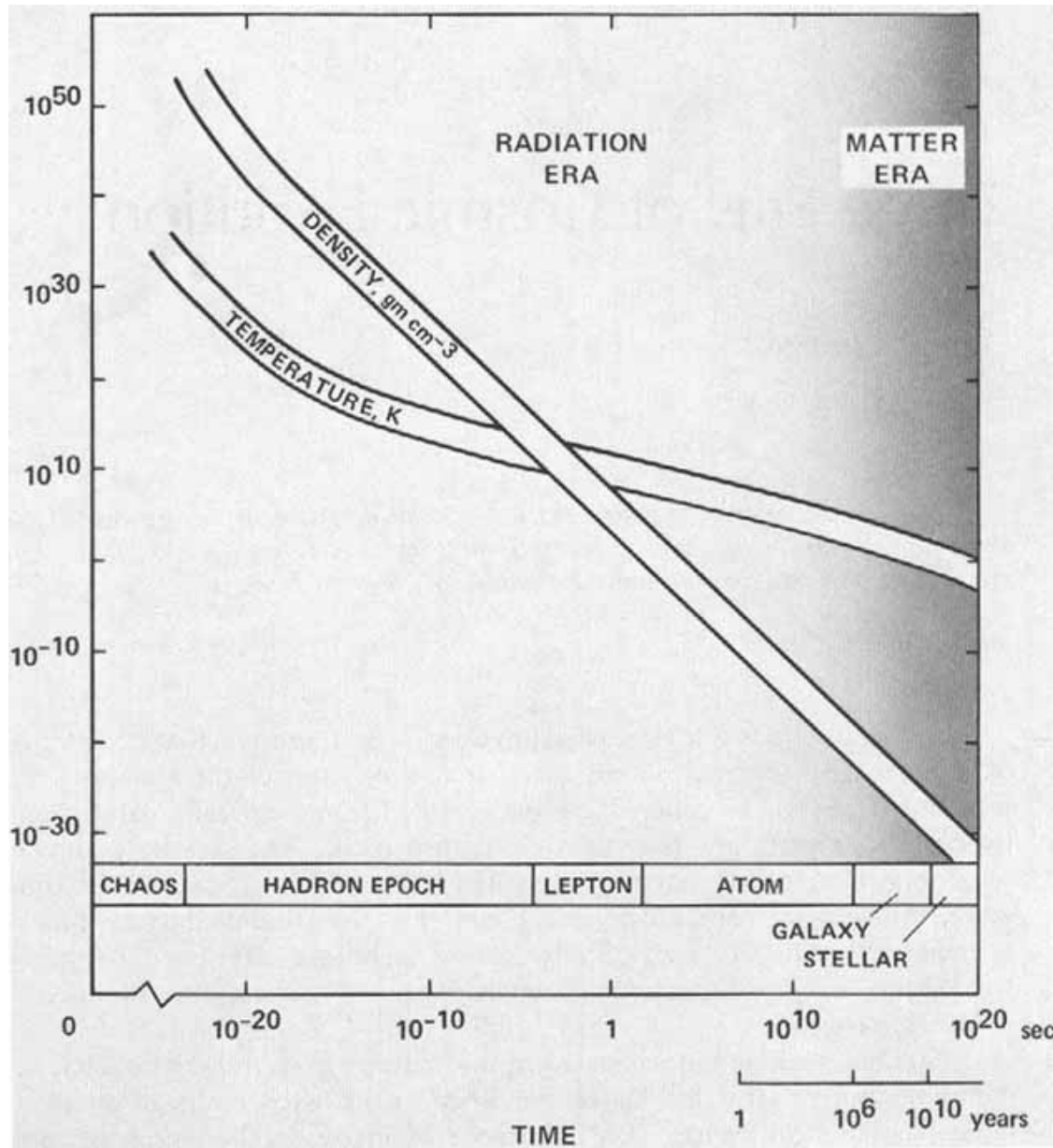


Big Bang

universe began at some point in time

→ based on extrapolating back
the observed Hubble expansion
of galaxies

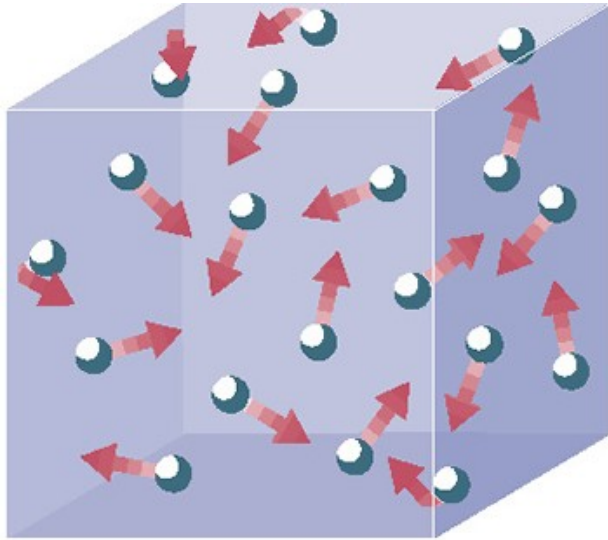
Is there any other evidence?



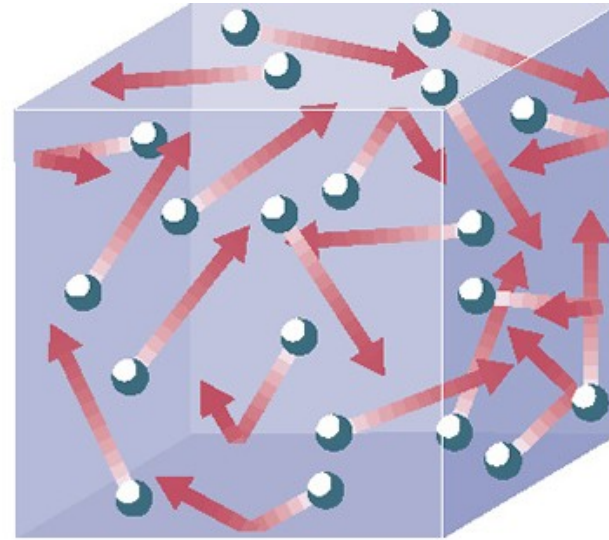
smaller in the past
 +
 roughly the same
 amount of matter
 and energy
 =
 the density of
 matter and energy
 must have been
 higher in the past

universe was hotter in the past

lower T



higher T



Longer arrows mean higher average speed.

- temperature is proportional to the average kinetic energy per molecule

$$E_K = \frac{1}{2}mv^2 = \frac{3}{2}kT$$

Boltzmann constant

$$k = 1,4 \cdot 10^{-23} \text{ J/K} = 8,6 \cdot 10^{-5} \text{ eV/K}$$

The Big Bang

15 thousand million years

1 thousand million years

300 thousand years

3 minutes

1 second

10^{-10} seconds

10^{-34} seconds

10^{-43} seconds

10^{32} degrees

10^{27} degrees

10^{15} degrees





10^{10} degrees

10^9 degrees

3000 degrees

18 degrees

3 degrees K

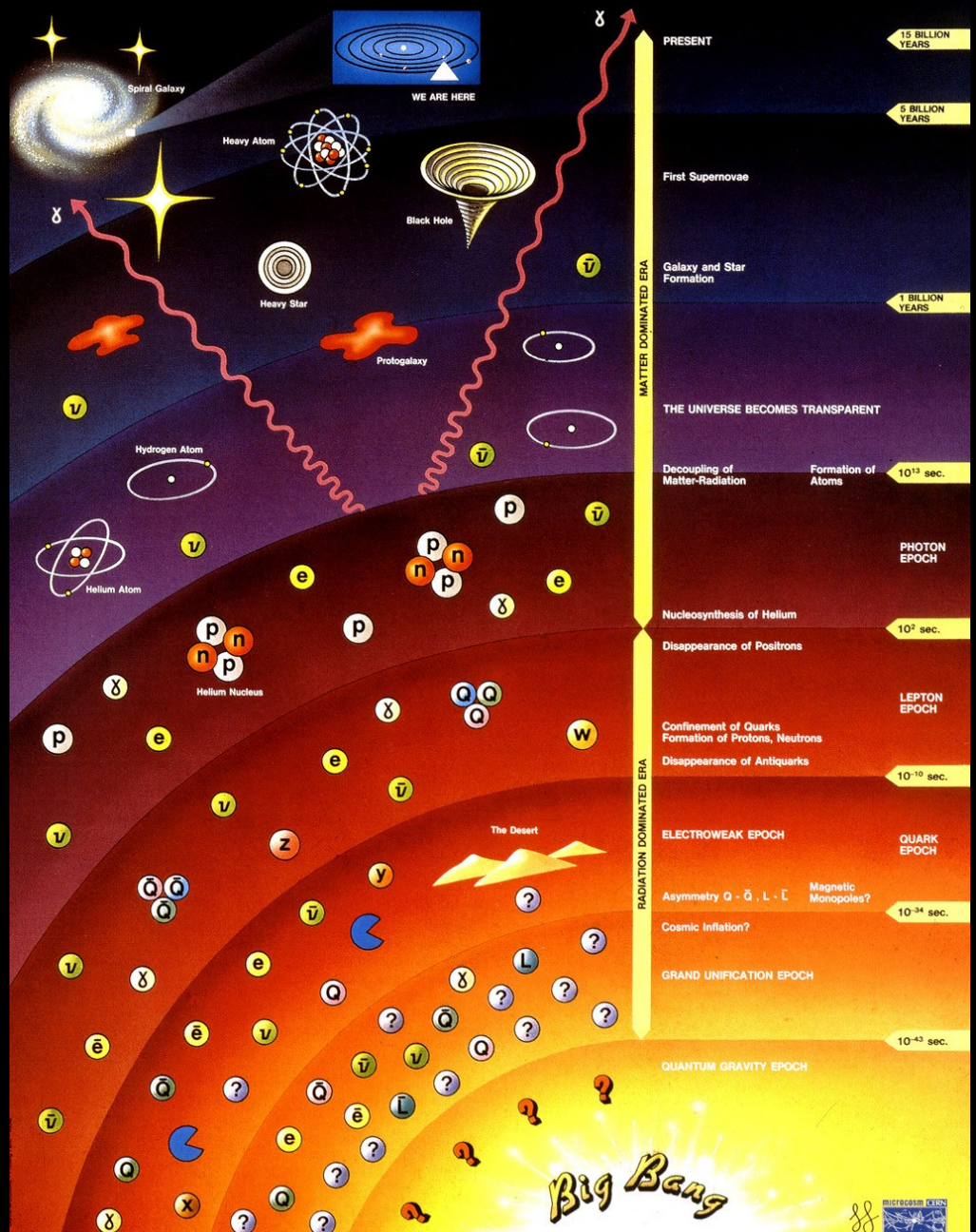
-  radiation
-  particles
- W^+ } heavy particles carrying the weak force
- W^- }
- Z }
-  quark
-  anti-quark
- e^- electron

-  positron (anti-electron)
-  proton
-  neutron
-  meson
- H hydrogen
- D deuterium
- He helium
- Li lithium



H. J. ...

History of the Universe



Big Bang



Disappearance of Antiquarks

10^{-10} sec.

The Desert

ELECTROWEAK EPOCH

QUARK EPOCH

Asymmetry $Q - \bar{Q}$, $L - \bar{L}$

Magnetic Monopoles?

10^{-34} sec.

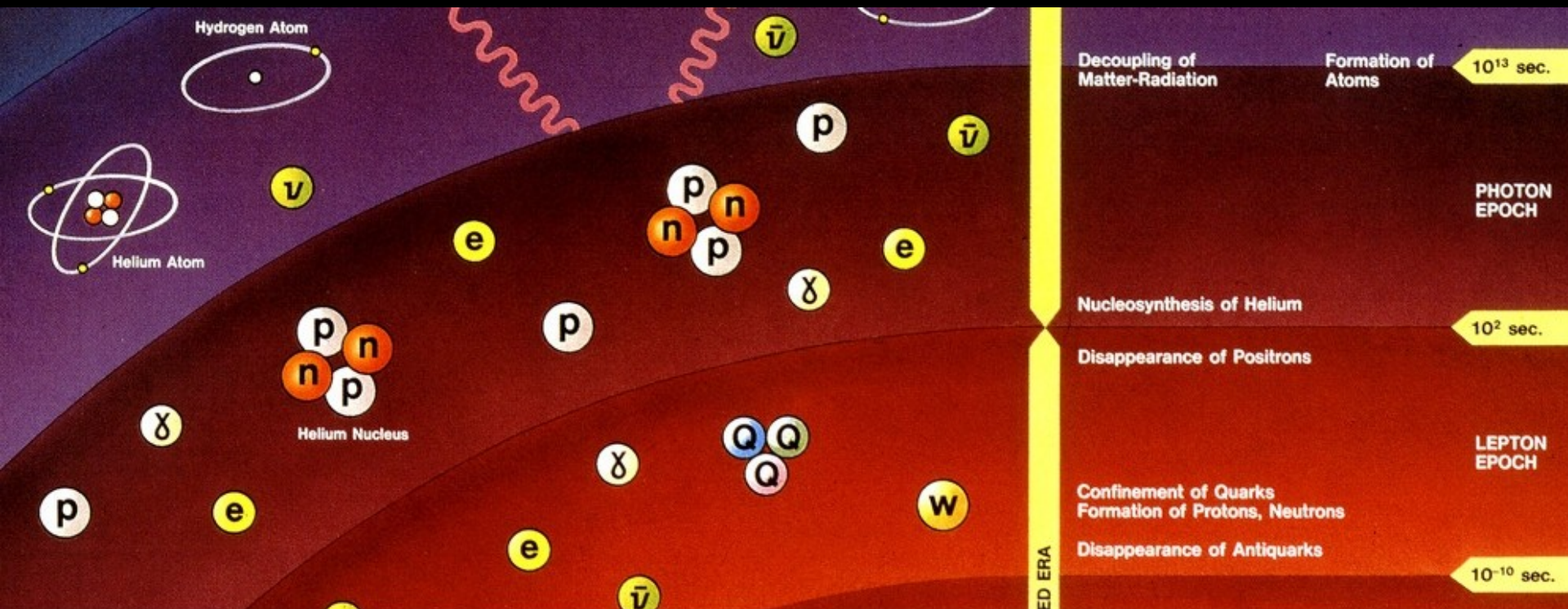
Cosmic Inflation?

GRAND UNIFICATION EPOCH

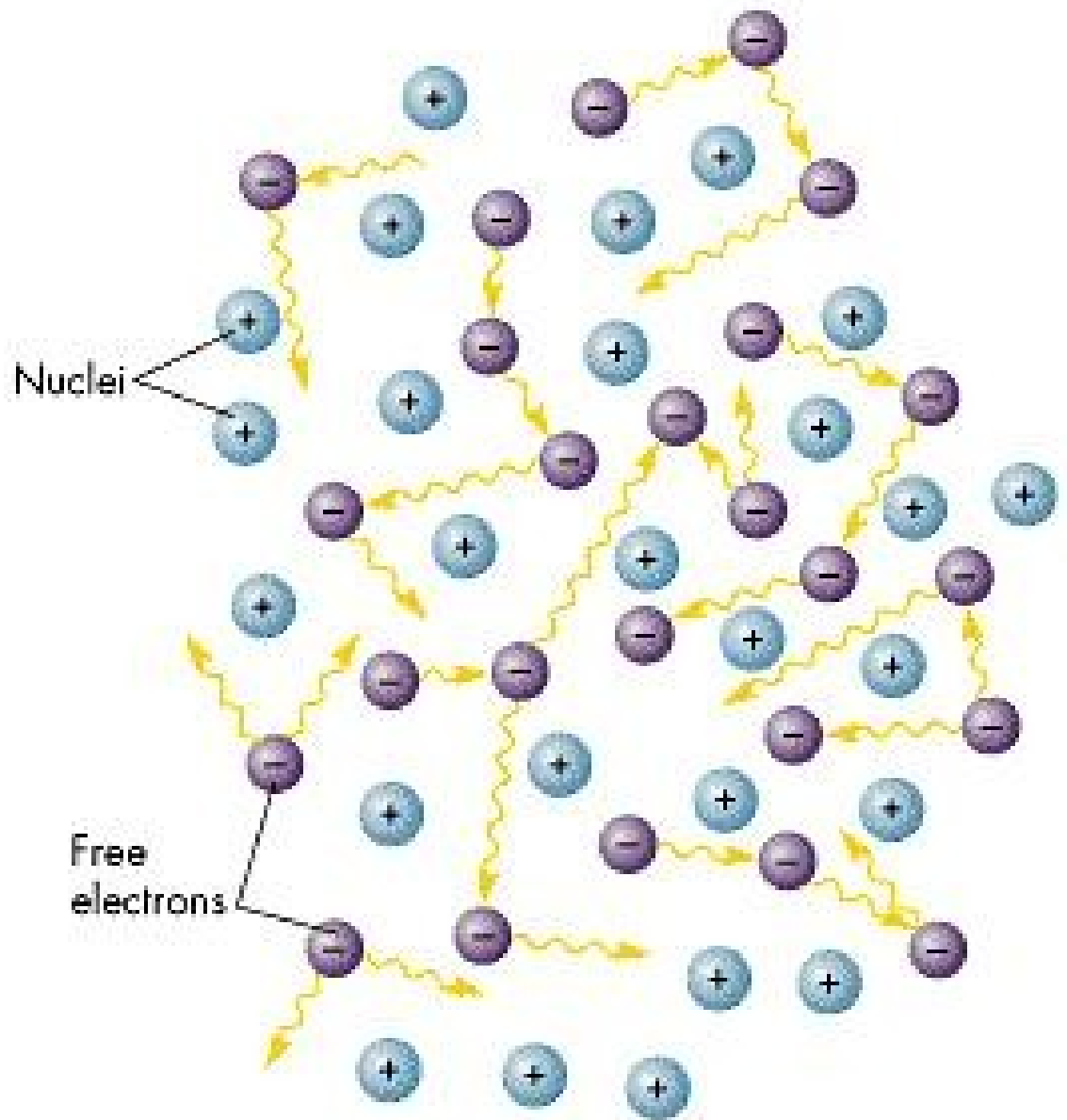
10^{-43} sec.

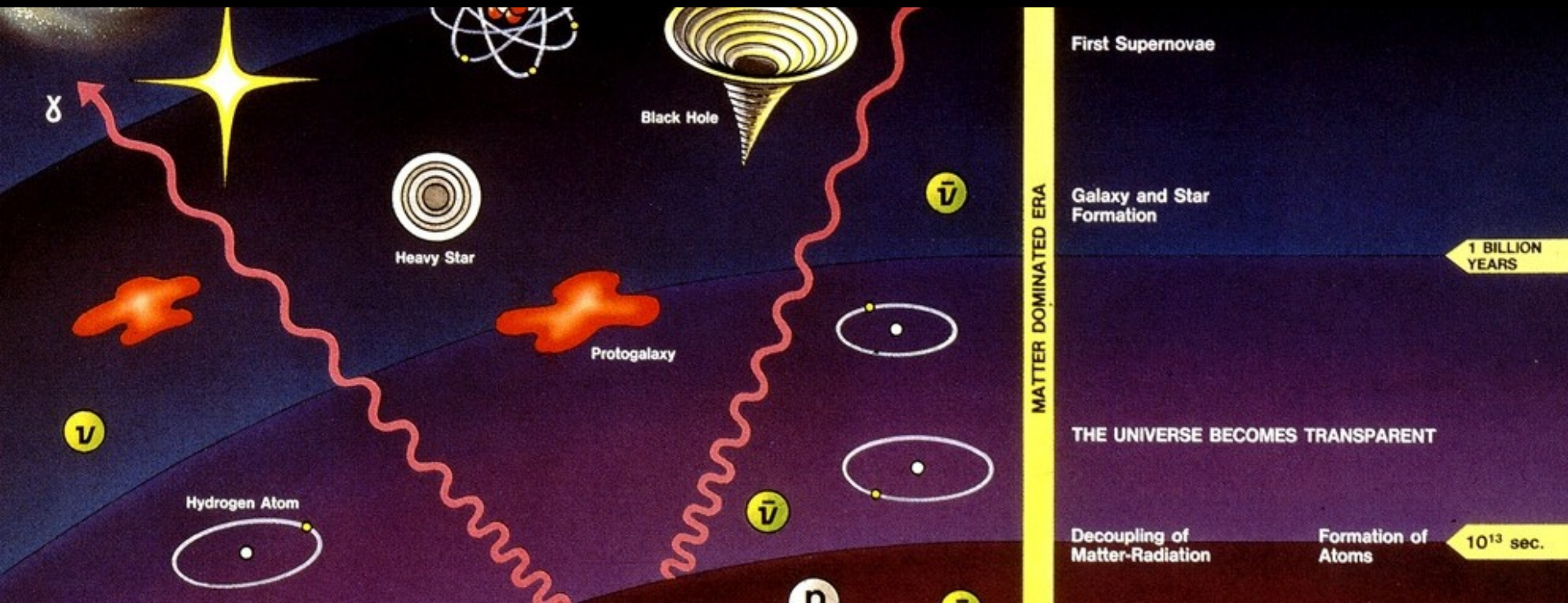
QUANTUM GRAVITY EPOCH

Big Bang

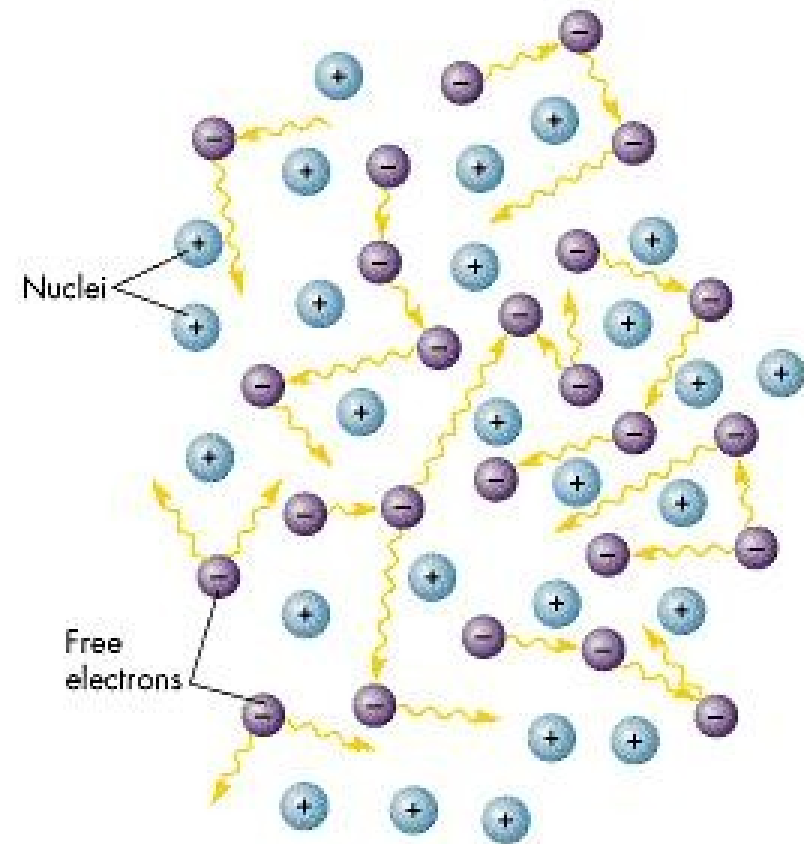


- first protons and neutrons at about 1 sec
- helium nuclei formed at about 100 sec
- observed ratio of helium/hydrogen matches Big Bang prediction
- universe is opaque

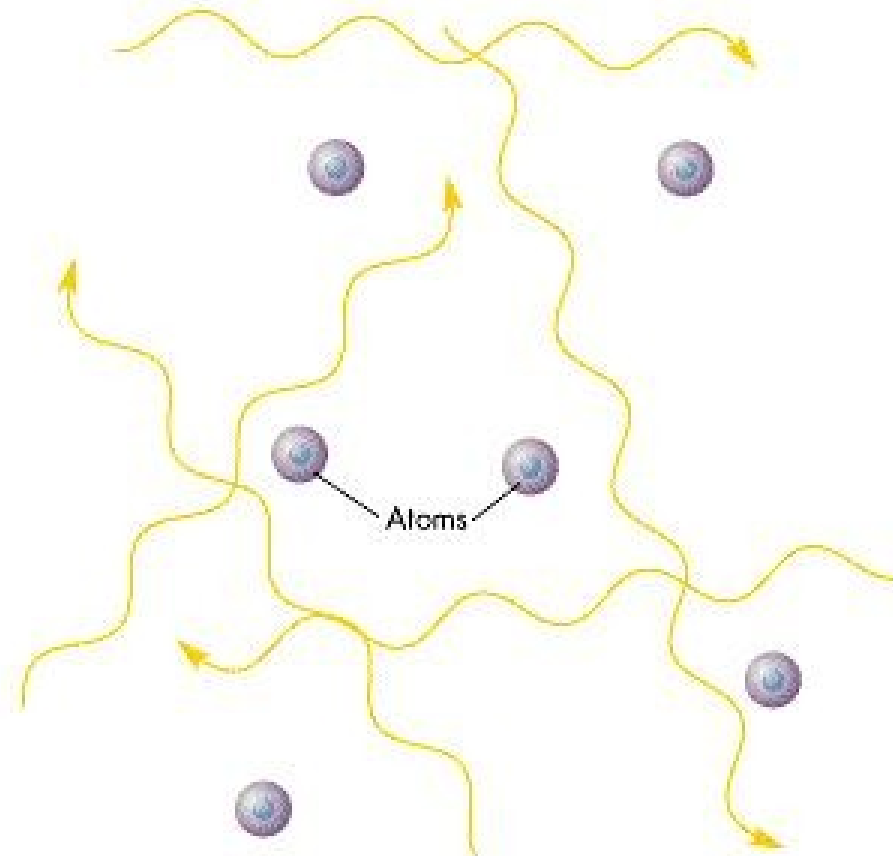




- at one million years, electrons combine with nuclei and atoms form
- universe becomes transparent



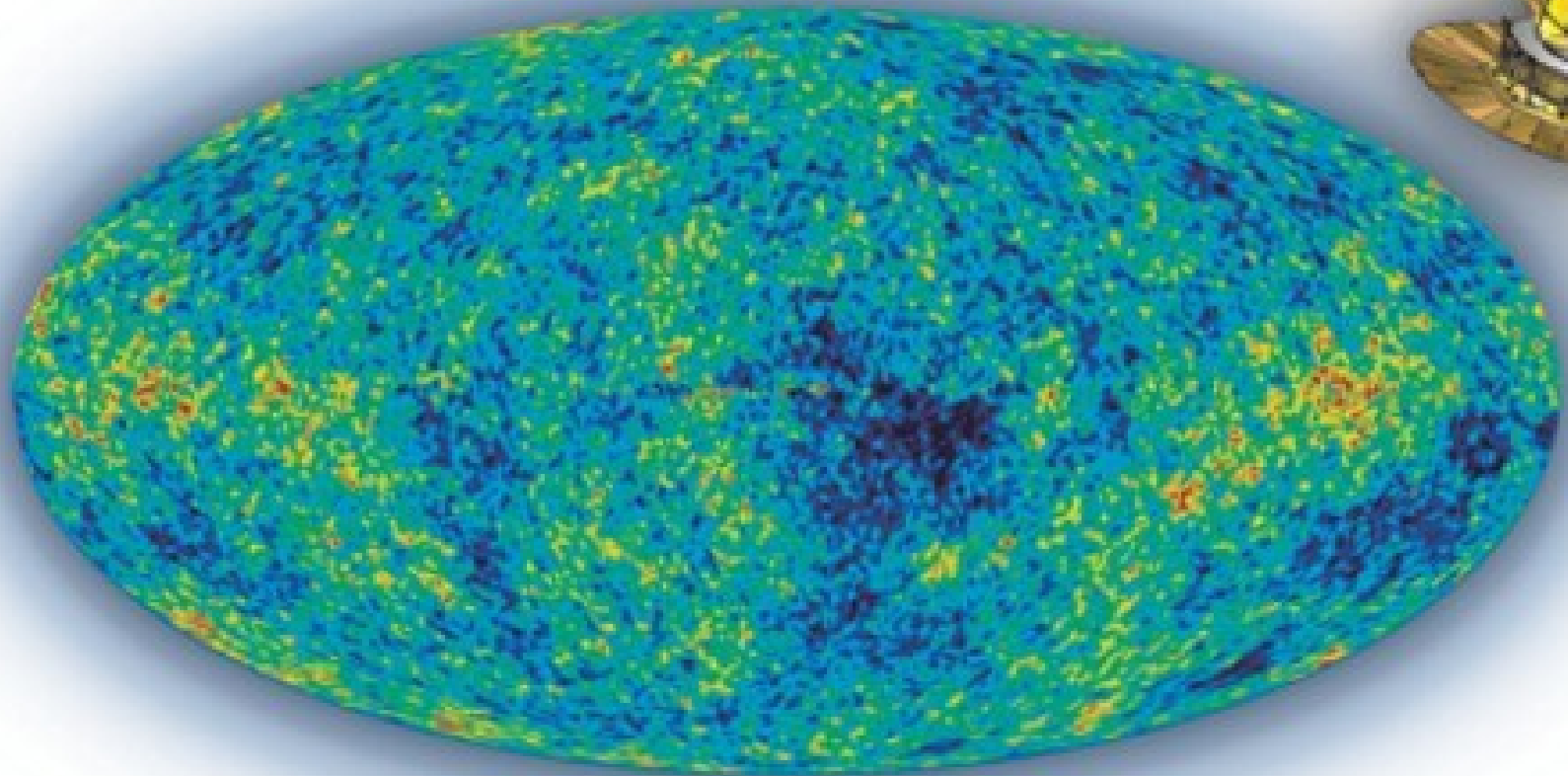
A Before recombination: The universe was opaque



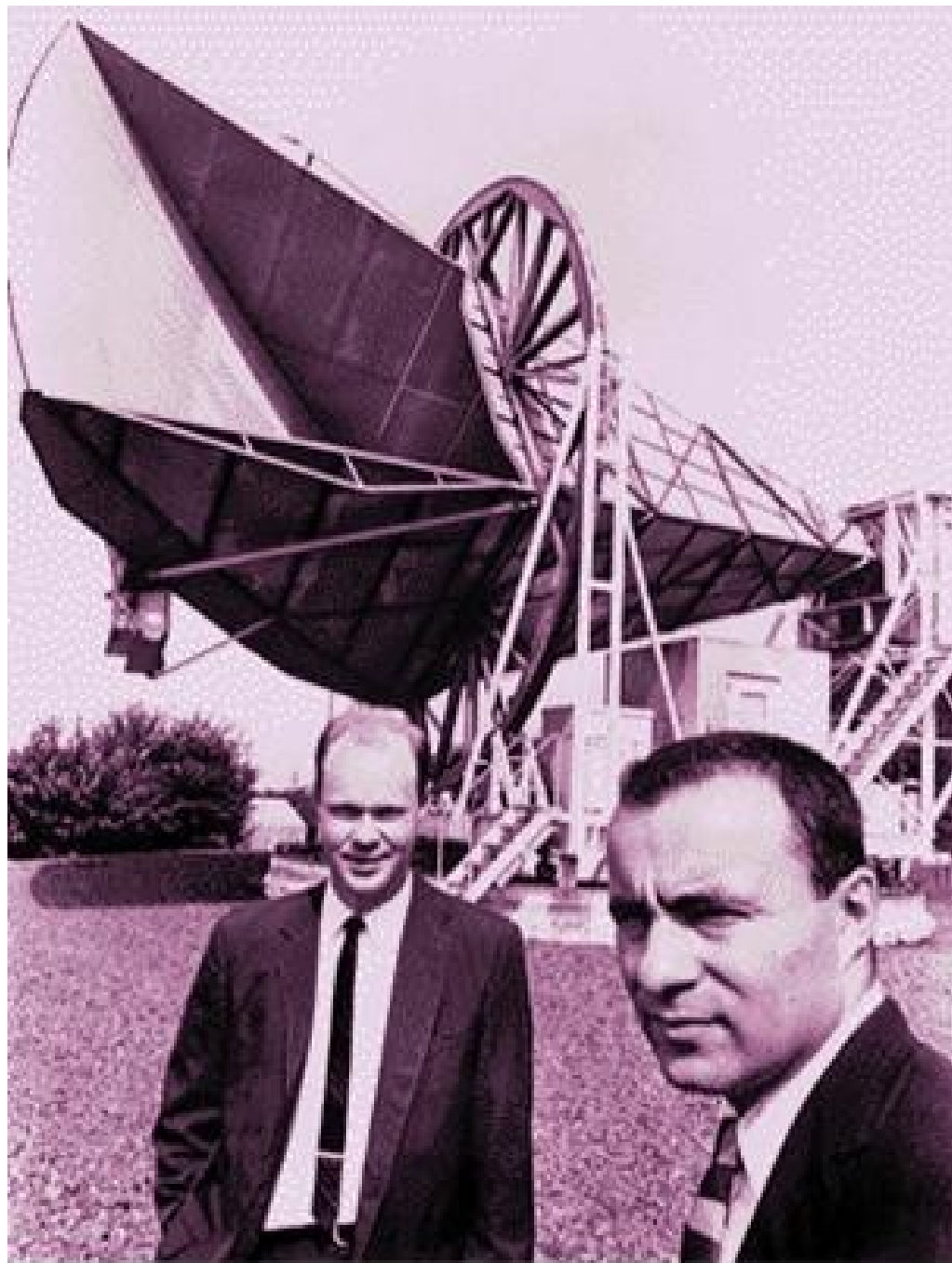
B After recombination: The universe was transparent

transition occurs at around $T = 3740$ K

cosmic microwave background (CMB)



universe glows at 2,7 K in every direction

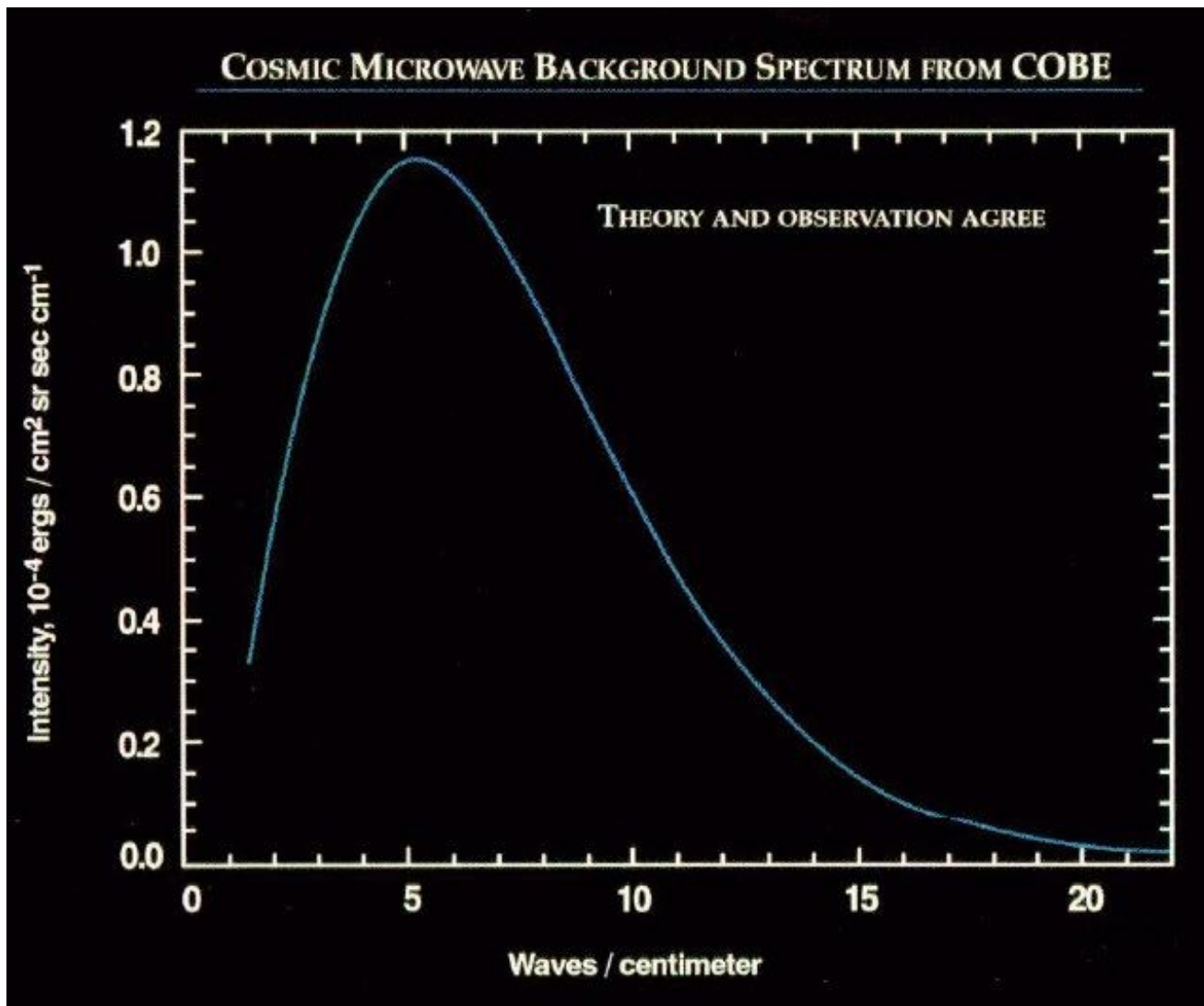


discovered by
Arno Penzias
and Robert Wilson

(1960-65)
AT&T's Bell Labs

finding the source of
noise in an antenna
used to bounce
telephone signals
bounced off metallic
balloons high in the
atmosphere

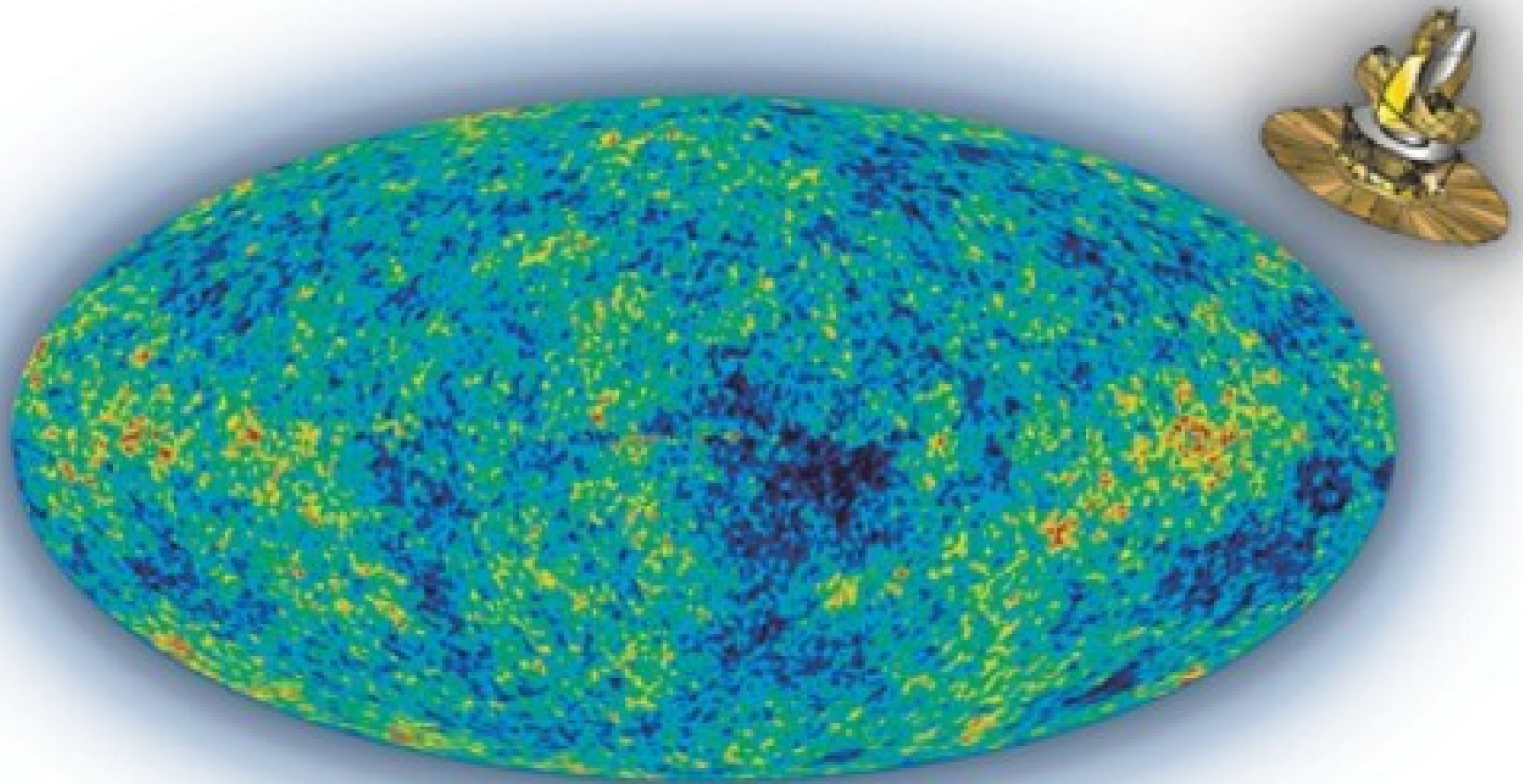
Nobel prize in 1978



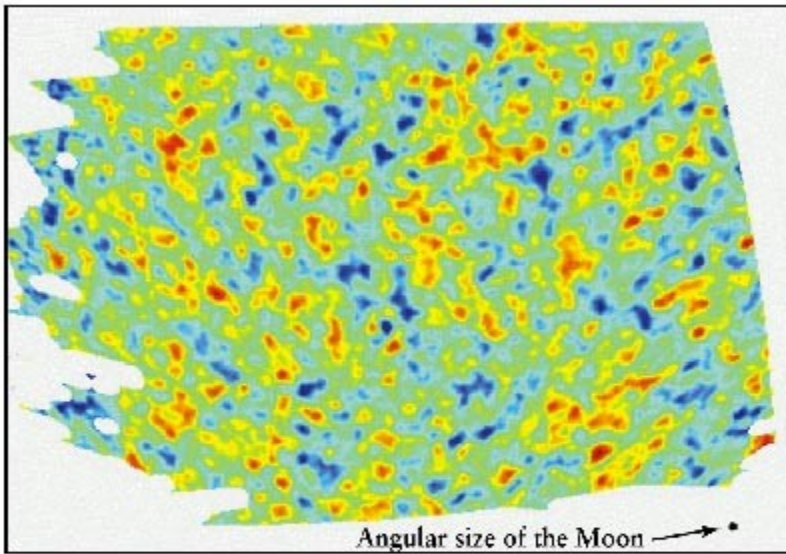
radiation is a blackbody spectrum emitted at 3000 K, red shifted by a factor of 1000

3 parts of evidence for the Big Bang model

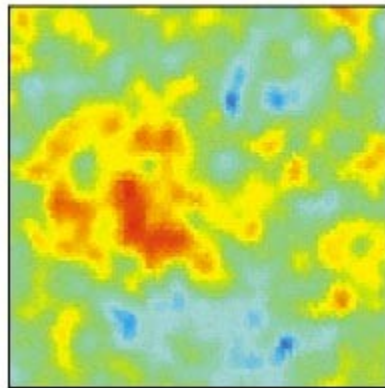
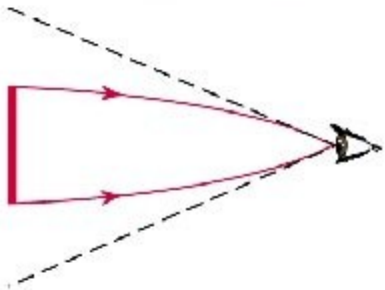
- Hubble expansion: galaxies are moving away from us with speed proportional to distance
- Ratio of Helium vs. Hydrogen in gas clouds unaffected by stars
- CMB:
a 2.7 K glow seen in all directions



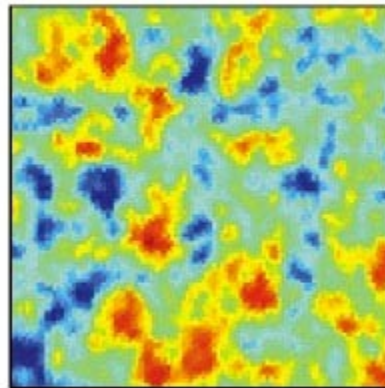
small fluctuations are due to sound waves
at recombination



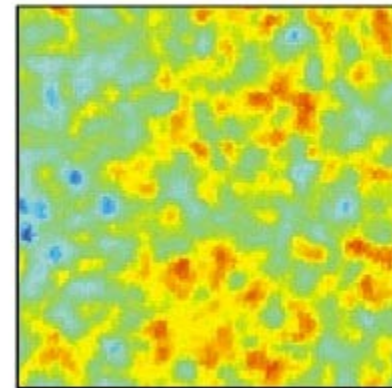
- temperature variations in the CMB are observed to be about 0.0003 K
- expected physical size of the hot/cold regions can be calculated



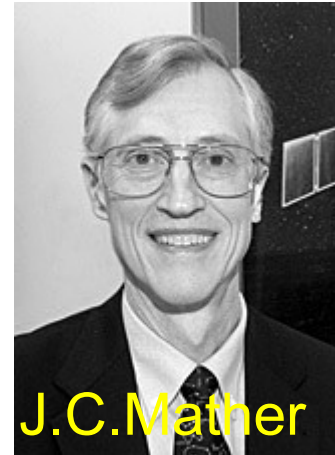
a If universe is closed, "hot spots" appear larger than actual size



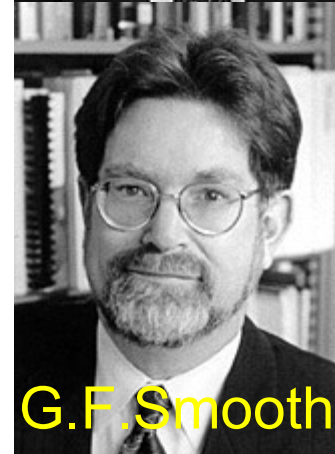
b If universe is flat, "hot spots" appear actual size



c If universe is open, "hot spots" appear smaller than actual size



J.C. Mather



G.F. Smooth

curvature of the universe

determined by the density parameter Ω_0

$$\Omega_0 = \frac{\rho}{\rho_c} \quad \Omega_0 < 1 \Rightarrow \text{negative curvature}$$
$$\Omega_0 > 1 \Rightarrow \text{positive curvature}$$

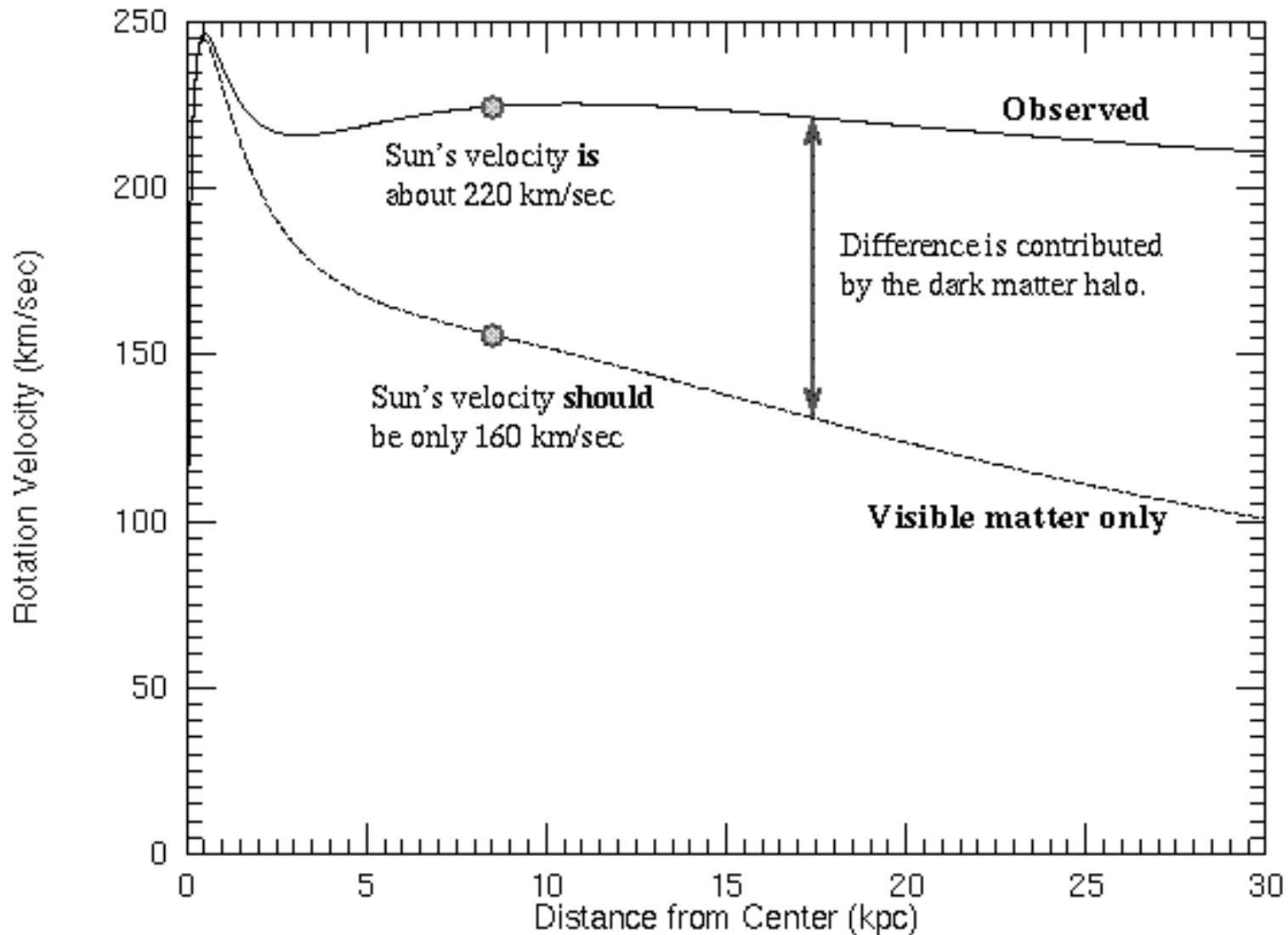
measurement of CMB fluctuations gives

$$\Omega_0 = 1.02 \pm 0.02$$

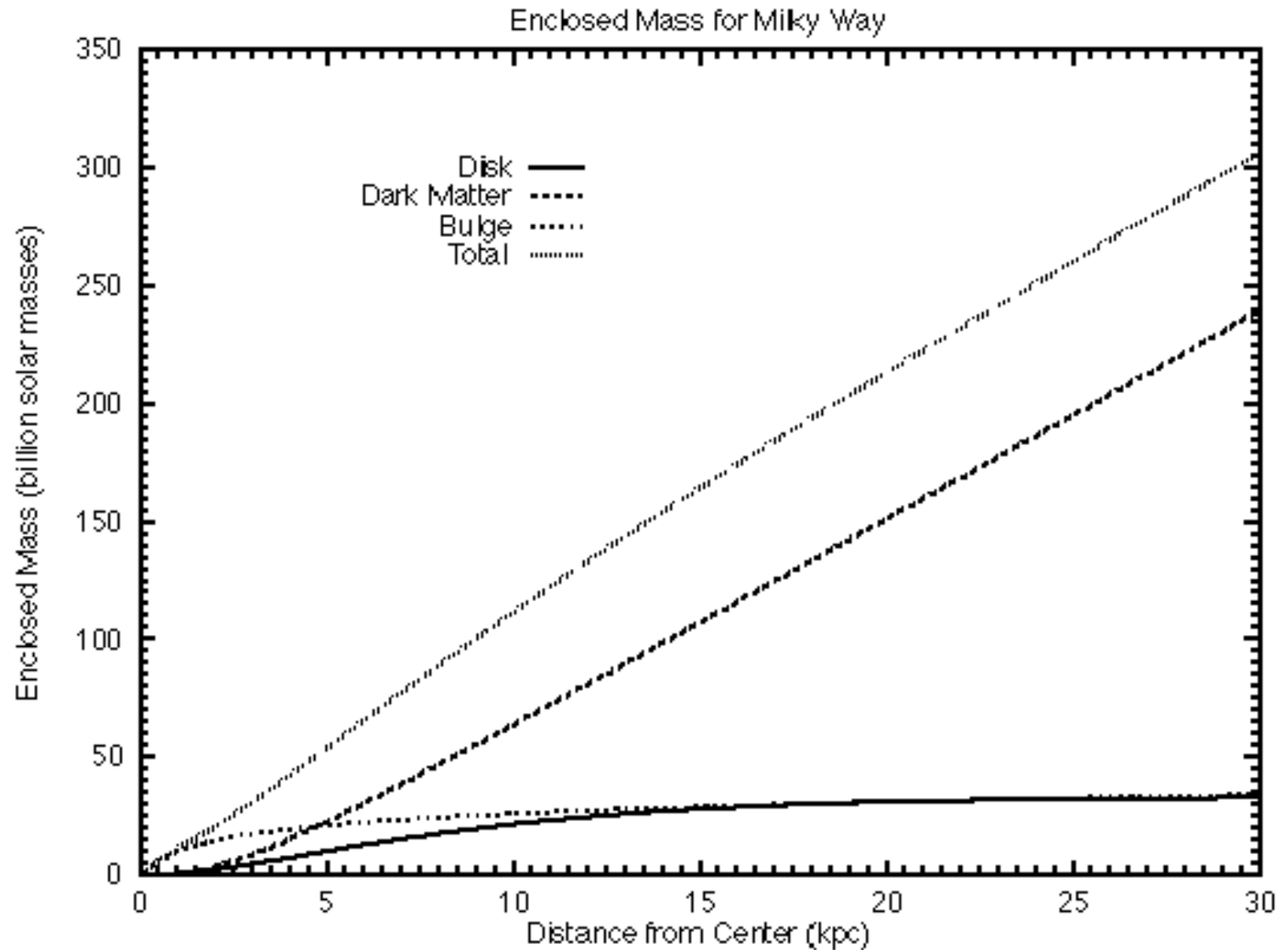
contents of the Universe

- normal matter
 - Stars
 - hot gas
 - anything made of atoms
- total is 4% of ρ_c

rotation curve of our Galaxy



mass of our Galaxy



dark matter

- dark – it doesn't produce light (of any kind)
- it does have mass, produces gravity
- nature is unknown

- most likely that's elementary particles

contents of the universe

- normal matter is 4% of ρ_C
- dark matter is 23% of ρ_C
- total of normal and dark matter is $\Omega_M = 0.3$
- but, we need 100% of ρ_C
- other 73% that is dark energy $\Omega_\Lambda = 0.7$

contents of the universe



cosmology

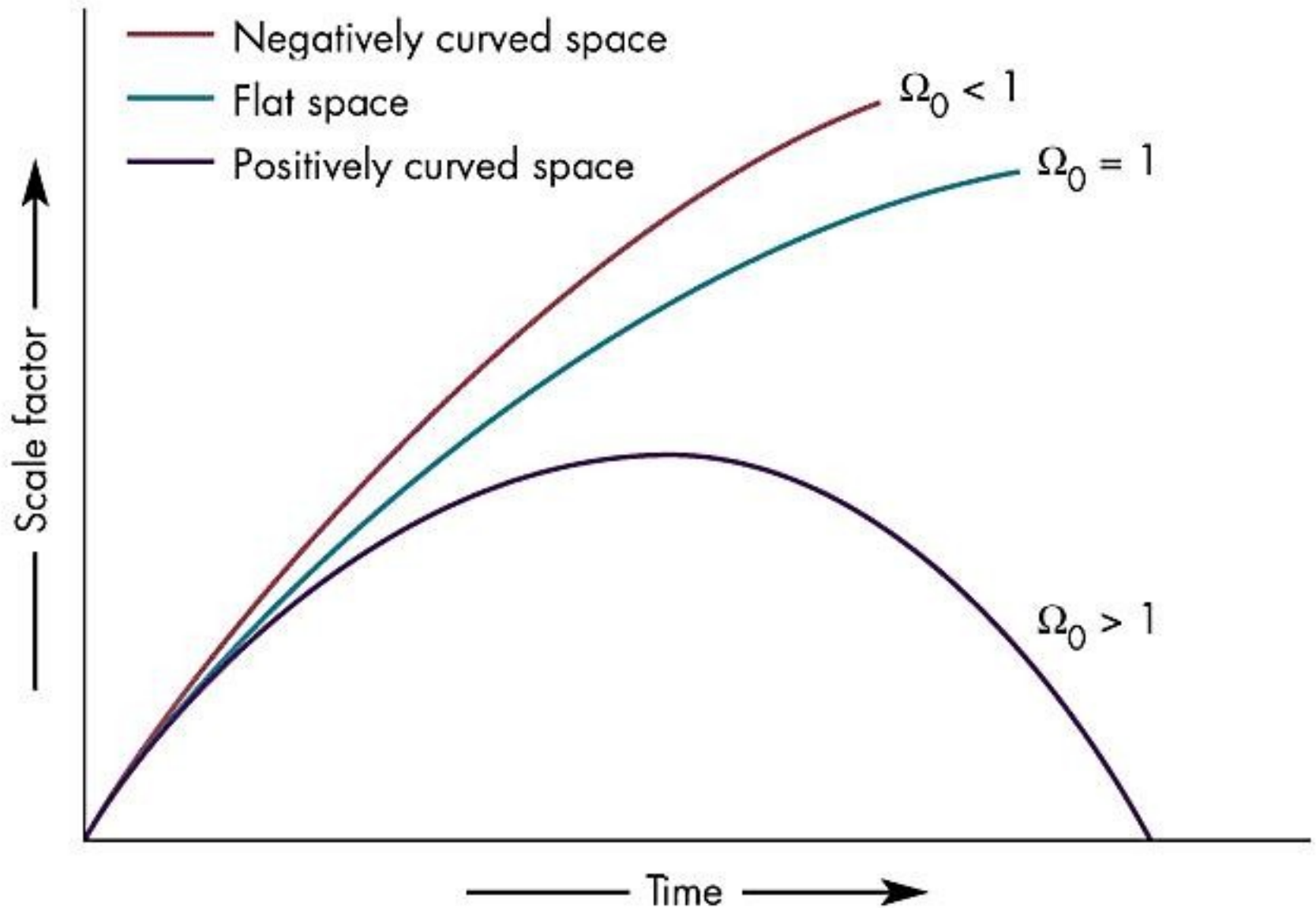
Einstein and cosmology

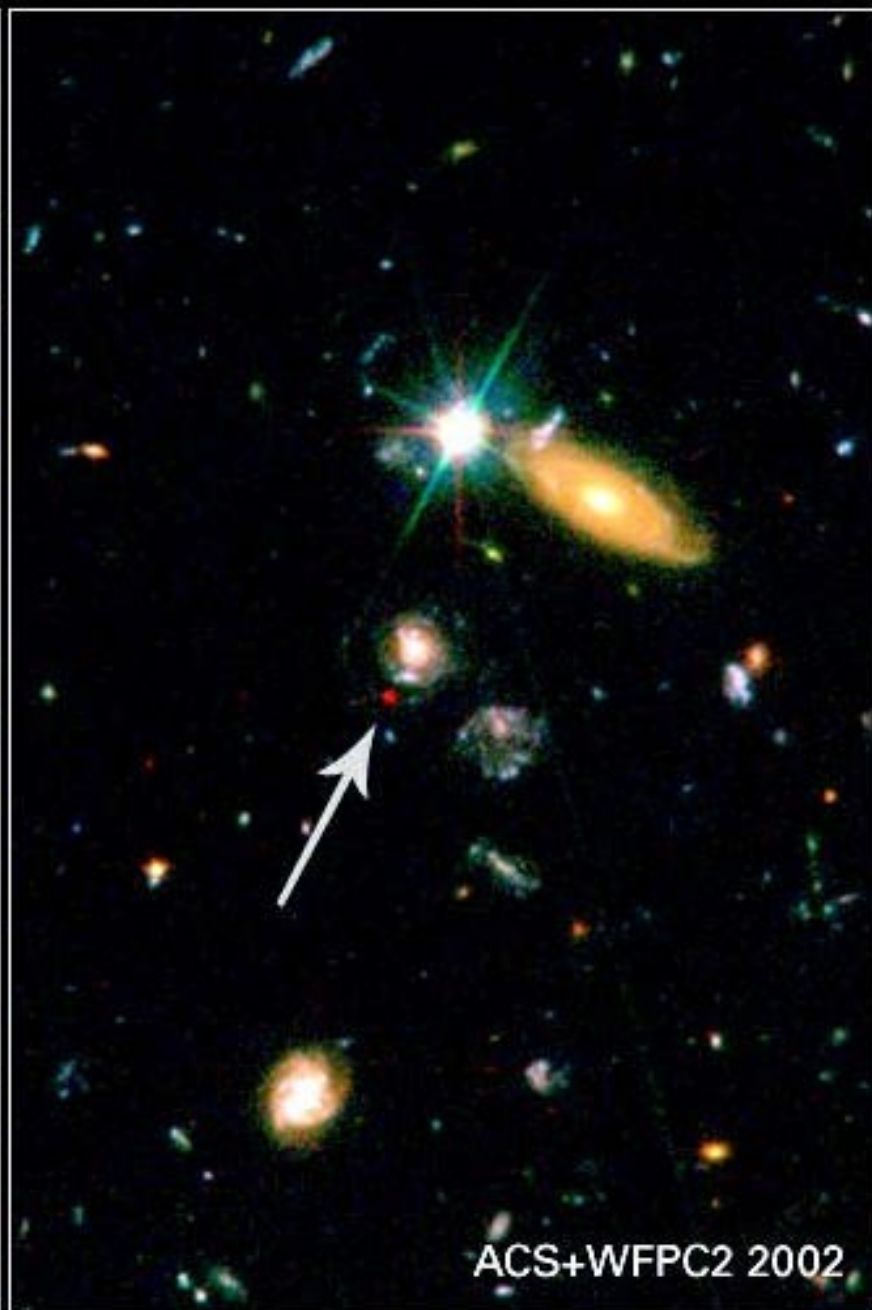
- Einstein:
 - wrote down the equations for GR
 - made a model of the universe
 - found that universe had to be either expanding or contracting
- new term (cosmological constant, Λ):
 - representing a energy field which could create antigravity to allow a static model
- after Hubble, Einstein called Λ his greatest blunder

cosmological constant

- quantum physics predicts that some energy fields that act like Λ
- one such field is the one thought to cause the rapid expansion of the Universe during inflation
- another such field appears to be operating today

matter slows down expansion

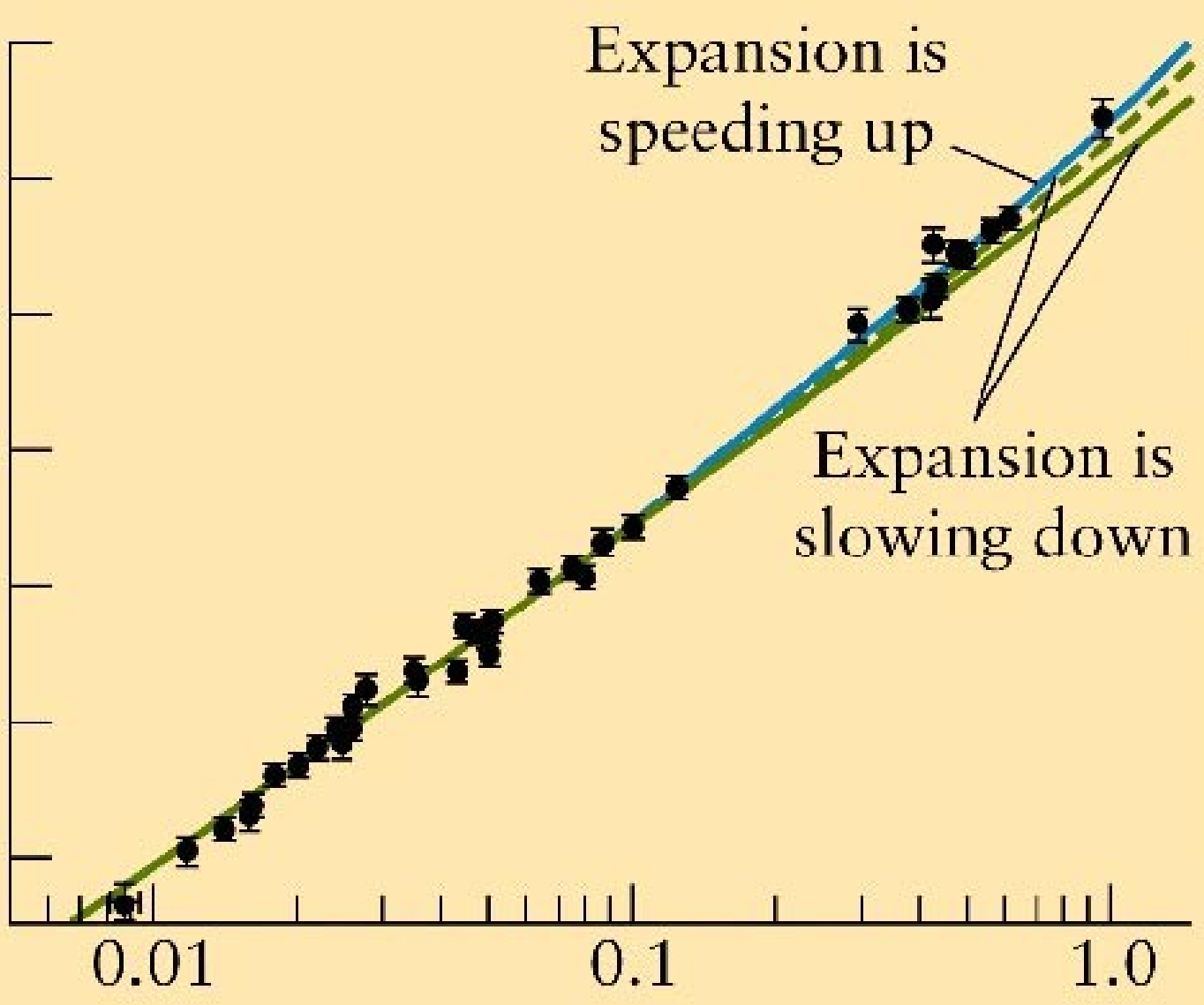




Distance \rightarrow
 \leftarrow Apparent brightness

Apparent magnitude of supernova

26
24
22
20
18
16
14

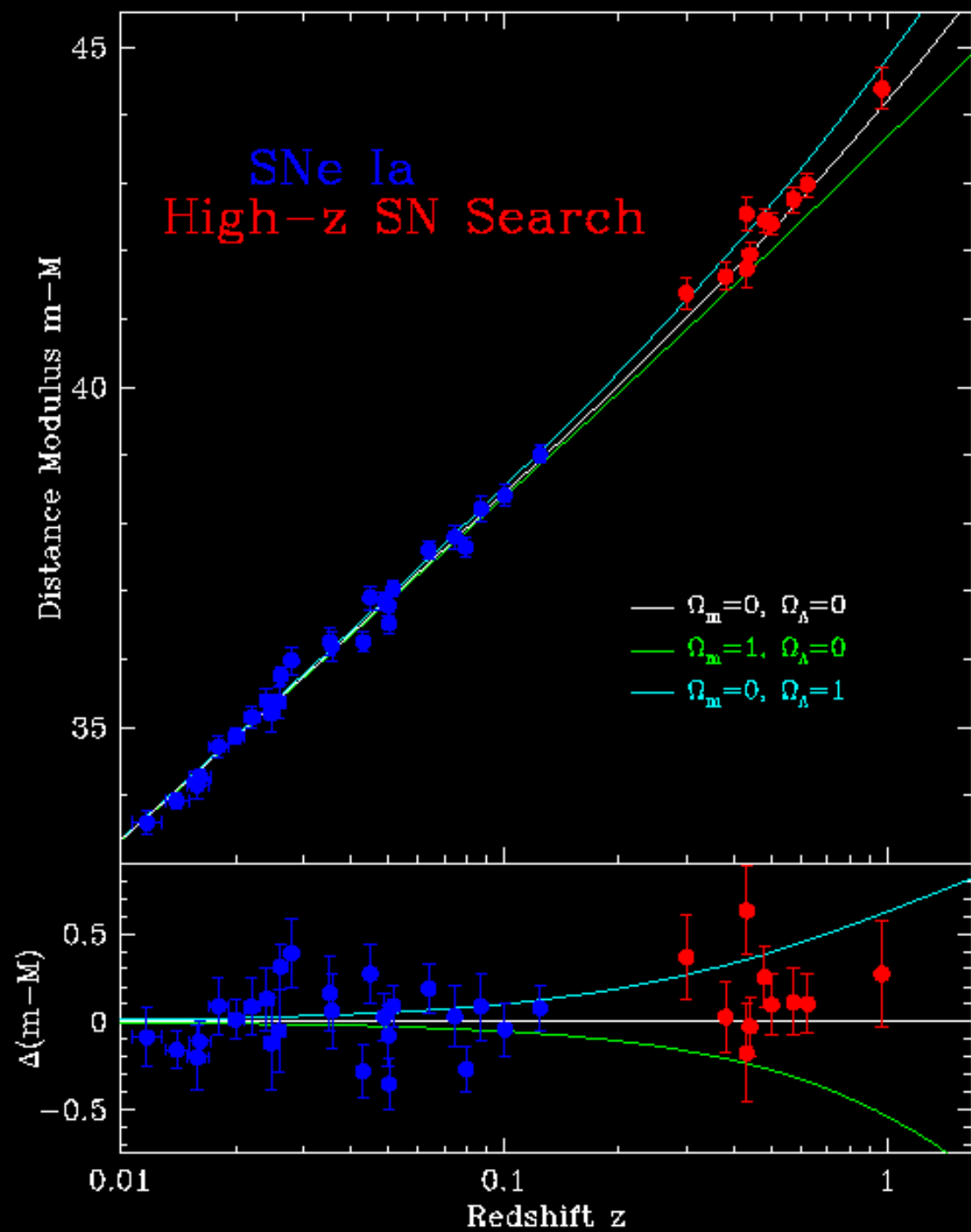


Expansion is speeding up

Expansion is slowing down

Redshift (z)

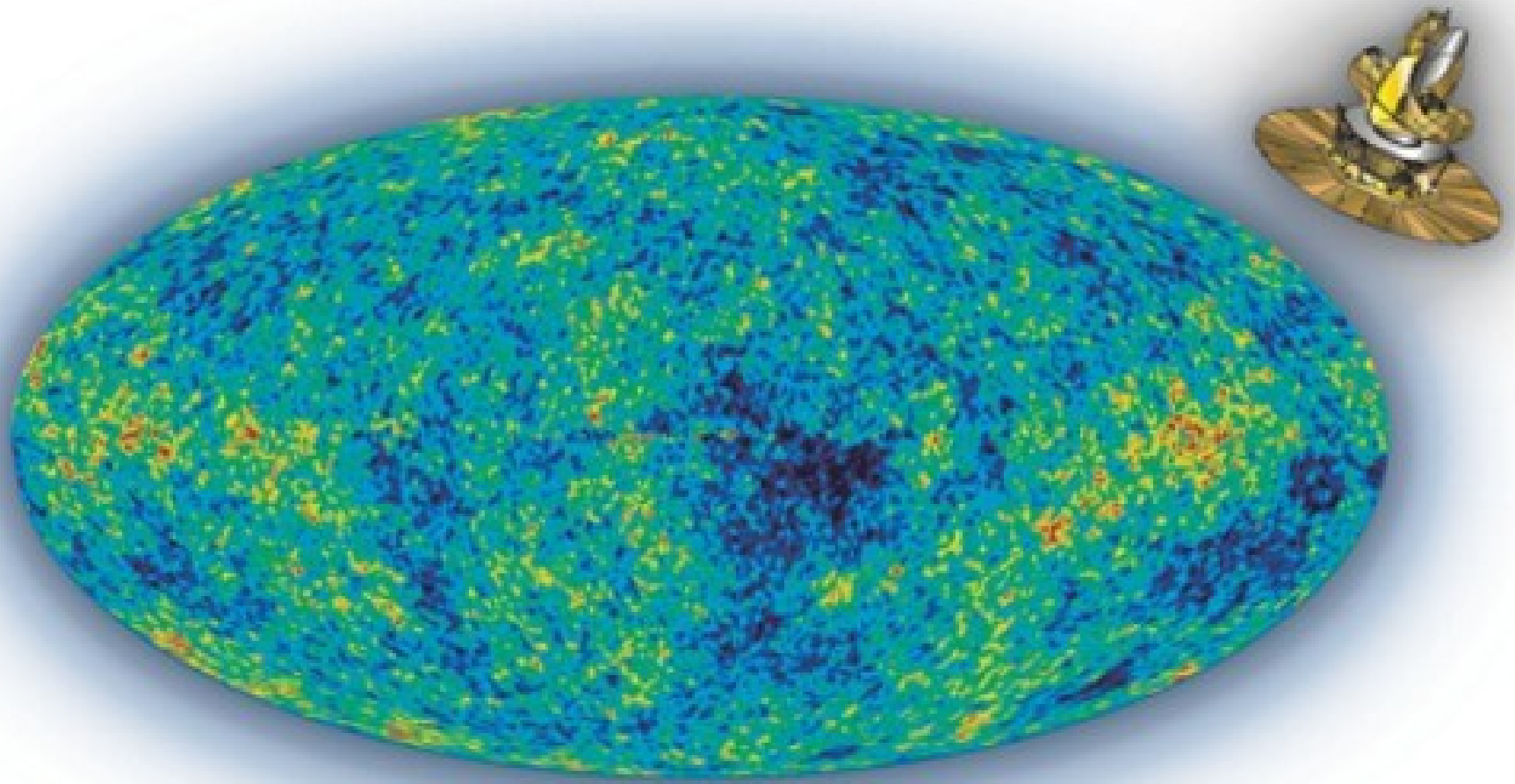
Recessional velocity \rightarrow



accelerating universe

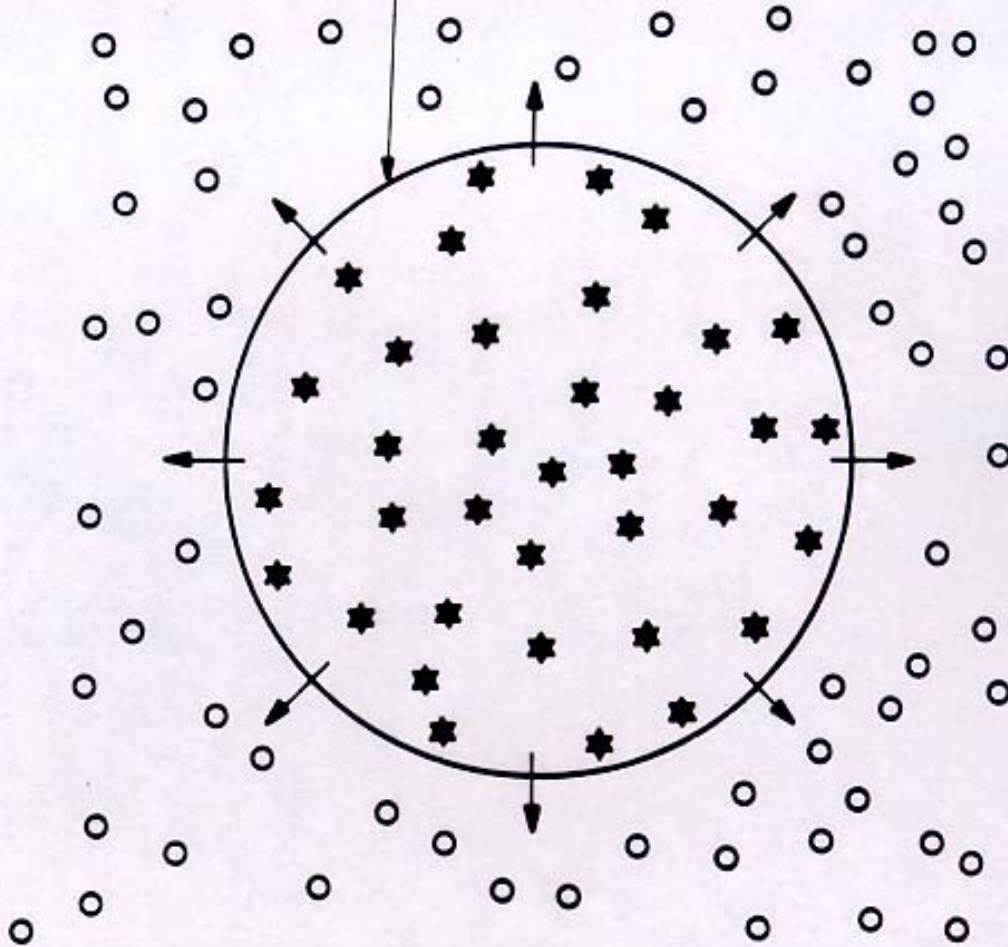
- Hubble expansion appears to be accelerating
- normal matter cannot cause acceleration, only deceleration of expansion
- dark energy is required
 - may be cosmological constant
 - may be something else
 - major current problem in astronomy

Problems with the Big Bang



universe glows at 2,7 K in every direction
temperature is the same to $< 0,1 \%$

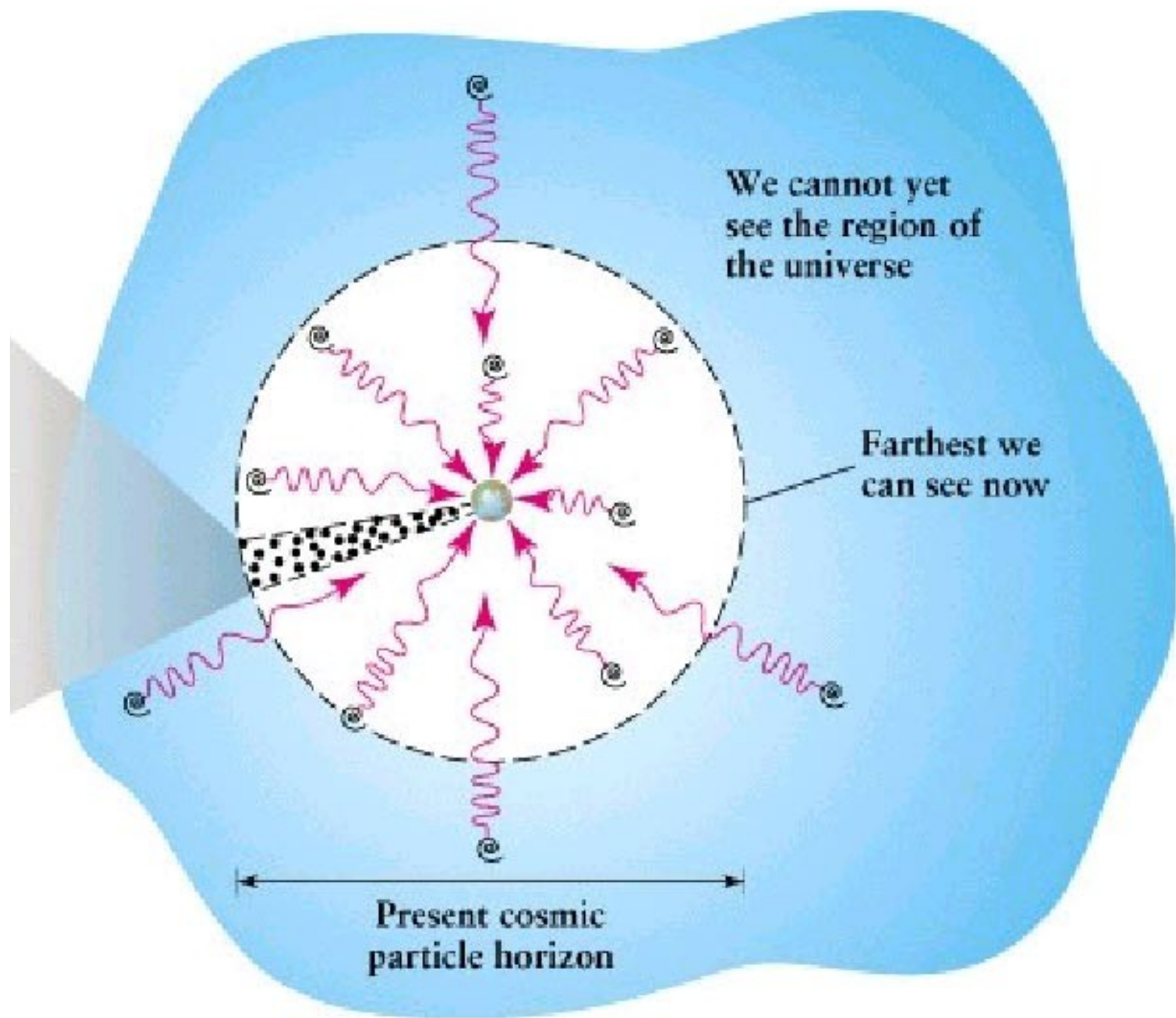
Edge of observable universe



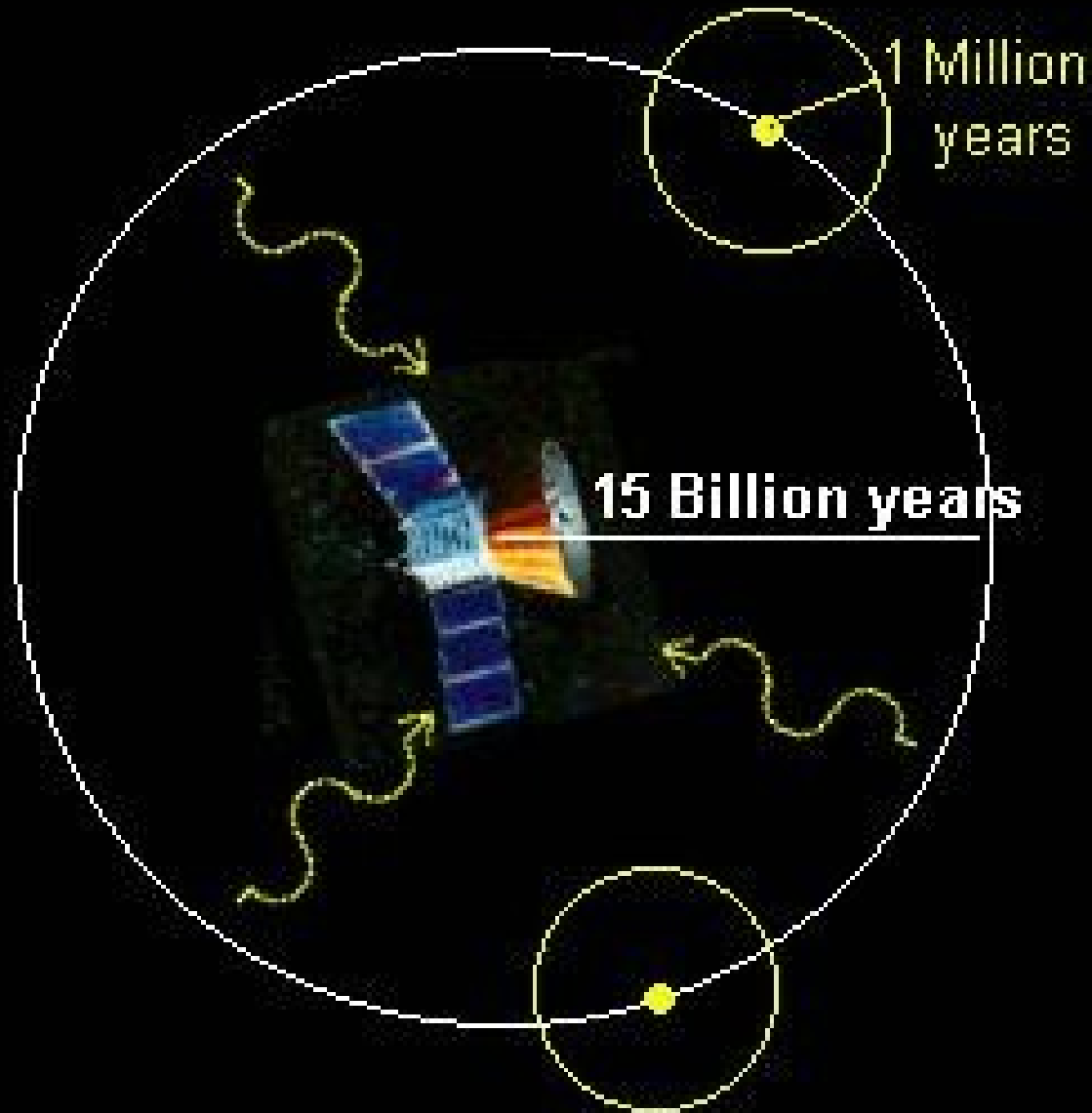
- ★ Stars visible
- Stars not yet visible

observable universe

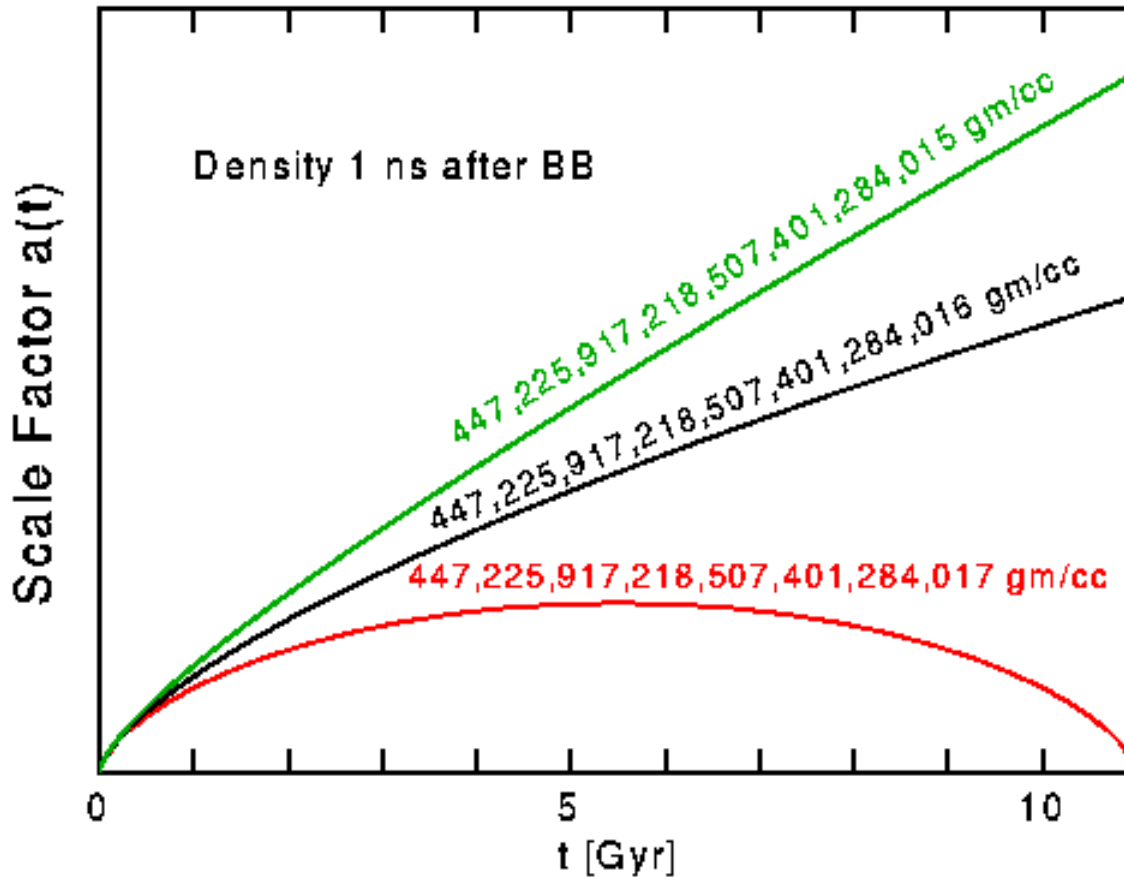
We can only see the parts of the Universe from which light has had time to travel to us!



horizon problem

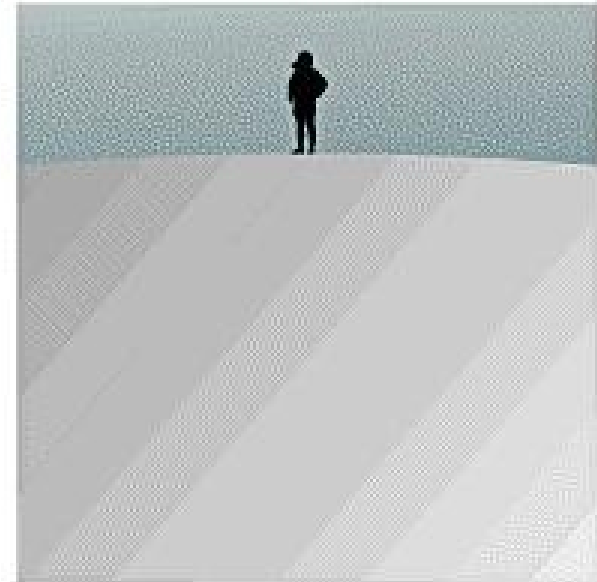
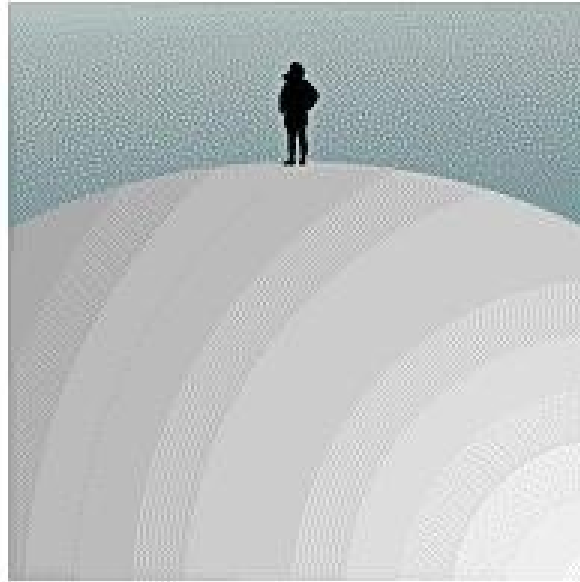
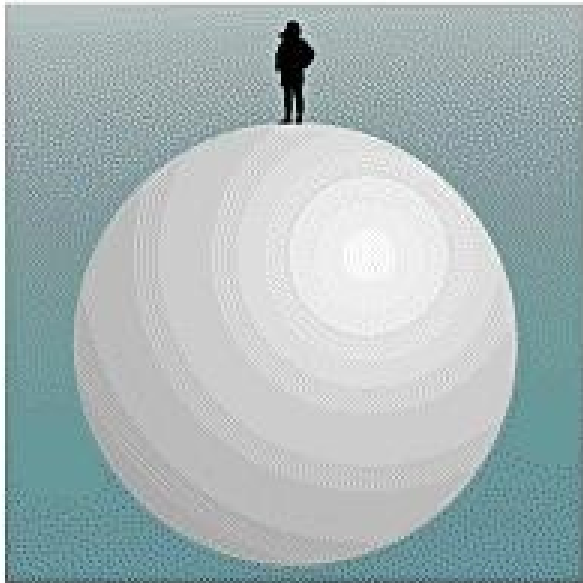


flatness problem

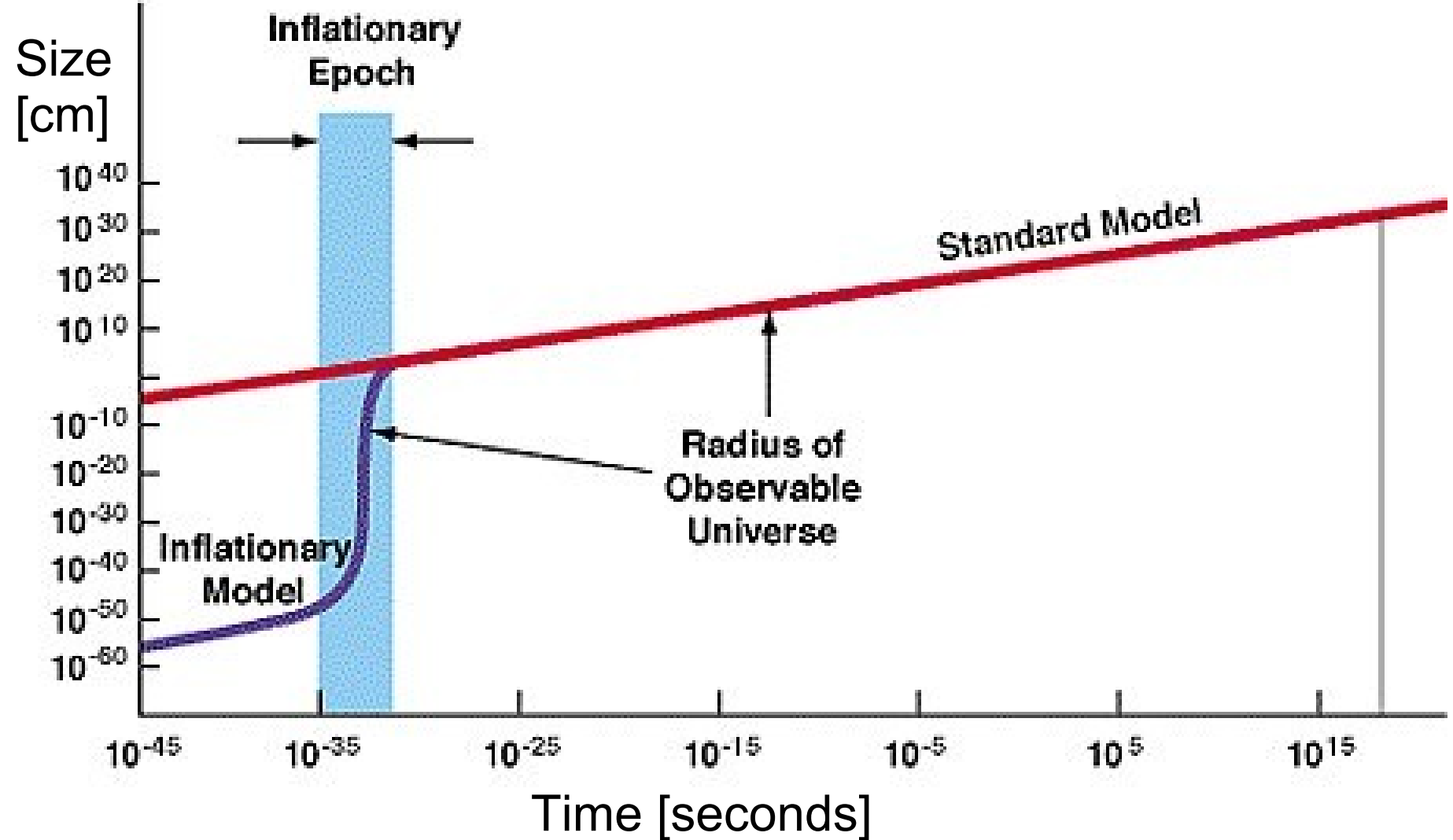


any tiny deviation from the critical density is amplified over time

inflation makes the Universe flat



inflation



whole observable universe came from a tiny region

inflation and cosmology

- universe is very uniformity and very close to flat
- new theoretical idea – inflation - gives an explanation for the uniformity and flatness of the universe
- inflation still needs to be tested
- NASA plans to fly a satellite to measure the polarization of the cosmic microwave background in order to test inflation